V&V Summary Report L2 ASCDS Version: 10.9.2

Observation 22997 - L2 Version 1 Chandra X-Ray Center

L2 Processing Date : Oct 16 2020

See axaff22997N001_VV001_vvref2.pdf for the full report

V&V Scientist	Melania Nynka
V&V Date (YYYY-MM-DD)	2020.10.19
V&V Edition	1
V&V Disposition and Status	OK
V&V Charge Time	26.943837353945

Comments

The focal plane temperature during the interval 719178083.62 - 719183885.22 (MET s) of this observation was warmer than the upper limit for optimum calibration of the ACIS gain and spectral resolution (i.e., -111.0 C for ACIS-S).

The Chandra calibration team calibrates the ACIS gain and spectral resolution using data from the external calibration source (ECS). ECS data show that the frontside-illuminated (FI) CCDs are more temperature sensitive than the backside-illuminated (BI) CCDs.

A summary of the current calibration status of the ACIS gain and spectral resolution can be found at:

http://asc.harvard.edu/cal/Acis/Cal_prods/Gain_and_Spectral_Resolution/ACIS_response_summary.html

The main points are:

- 1) The gain on BI chips remains within 0.3% (i.e., the systematic uncertainty in the ACIS gain quoted on the Chandra Calibration Status Summary web page) at all measured temperatures.
- 2) The gain on FI chips remains within 0.3% below row 600 at all measured temperatures.

- 3) The gain on FI chips above row 600 can be underestimated by as much as 1% for focal plane temperatures exceeding -116 C.
- 4) The spectral resolution (i.e., FWHM) on BI chips is insensitive to the focal plane temperature.
- 5) Warmer focal plane temperatures increase the FWHM on FI chips by up to 30 eV near row 512 and by up to 70 eV near the top of the chips.

In summary, the user should be cautious in the spectral analysis of high S/N emission lines detected on the top half of FI chips in this observation. Default processing with the current version of the CALDB will underestimate photon energies by up to 1% and broaden emission lines by up to 70~eV.

201200	C 1
	Sequence number
22997	Observation id
THE TRUE NATURE OF X-RAYS FROM THE ORION TRAPEZIUM	Proposal title
Norbert Schulz	Principal investigator
Orion Nebula Cluster	Source name
0	& #160
P	events from which exps? Prim/Second/Both
83.81875	Observer's specified target RA [deg]
-5.38975	Observer's specified target Dec [deg]
83.820042493585	Nominal RA [deg]
-5.3920504085564	Nominal Dec [deg]
77.155489433506	Nominal Roll [deg]
1	Processing version of data
26943.837353945	Sum of GTIs [s]
26602.658261738	Livetime [s]
26943.878393888	Sum of GTIs [s]
26943.796313882	Sum of GTIs [s]
26943.755273938	Sum of GTIs [s]
26943.837353945	Sum of GTIs [s]
26943.714233875	Sum of GTIs [s]
26943.673193932	Sum of GTIs [s]
428159	Number of level 2 events
	THE TRUE NATURE OF X-RAYS FROM THE ORION TRAPEZIUM Norbert Schulz Orion Nebula Cluster 0 P 83.81875 -5.38975 83.820042493585 -5.3920504085564 77.155489433506 1 26943.837353945 26602.658261738 26943.878393888 26943.796313882 26943.755273938 26943.714233875 26943.673193932

