V&V Summary Report L2 ASCDS Version: 10.8

Observation 21549 - L2 Version 1 Chandra X-Ray Center

L2 Processing Date: Aug 13 2019

See axaff21549N001_VV002_vvref2.pdf for the full report

V&V Scientist	Beth Sundheim
V&V Date (YYYY-MM-DD)	2020.04.02
V&V Edition	2
V&V Disposition and Status	OK
V&V Charge Time	53.809718253732

Comments

One optional chip was dropped.

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The focal plane temperature during part of this observation was warmer than the upper limit for optimum calibration of the ACIS gain and spectral resolution (i.e., -112.0 C for ACIS-I).

The Chandra calibration team calibrates the ACIS gain and spectral resolution using data from the external calibration source (ECS). ECS data show that the frontside-illuminated (FI) CCDs are more temperature sensitive than the backside-illuminated (BI) CCDs.

A summary of the current calibration status of the ACIS gain and spectral resolution can be found at:

http://asc.harvard.edu/cal/Acis/Cal_prods/Gain_and_Spectral_Resolution/ACIS_response_summary.html

The main points are:

- 1) The gain on BI chips remains within 0.3% (i.e., the systematic uncertainty in the ACIS gain quoted on the Chandra Calibration Status Summary web page) at all measured temperatures.
- 2) The gain on FI chips remains within 0.3% below row 600 at all measured temperatures.

- 3) The gain on FI chips above row 600 can be underestimated by as much as 1% for focal plane temperatures exceeding -116 C.
- 4) The spectral resolution (i.e., FWHM) on BI chips is insensitive to the focal plane temperature.
- 5) Warmer focal plane temperatures increase the FWHM on FI chips by up to 30 eV near row 512 and by up to 70 eV near the top of the chips. In summary, the user should be cautious in the spectral analysis of high S/N emission lines detected on the top half of FI chips in this observation. Default processing with the current version of the CALDB will underestimate photon energies by up to 1% and broaden emission lines by up to 70 eV.

	1	
seq_num	801831	Sequence number
obs_id	21549	Observation id
title	Studying the Progenitors of Our Favorite Clusters at $z > 1$	Proposa
observer	Michael McDonald	Principal investigator
object	SPT-CLJ2306-5120	Source name
dtycycle	0	& #160
cycle	P	events from which exps? Prim/Second/Both
ra_targ	346.615	Observer's specified target RA [deg]
dec_targ	-51.347	Observer's specified target Dec [deg]
ra_nom	346.66515596694	Nominal RA [deg]
dec_nom	-51.342301258641	Nominal Dec [deg]
roll_nom	39.247863122553	Nominal Roll [deg]
revision	1	Processing version of data
ontime	53809.718253732	Sum of GTIs [s]
livetime	53106.654670608	Livetime [s]
ontime0	53812.859333754	Sum of GTIs [s]
ontime1	53809.718323469	Sum of GTIs [s]
ontime2	53809.718293548	Sum of GTIs [s]
ontime3	53809.718253732	Sum of GTIs [s]
12events	150897	Number of level 2 events

