## V&V Summary Report L2 ASCDS Version : 10.6

## Observation 20828 - L2 Version 2 Chandra X-Ray Center

L2 Processing Date : Oct 31 2017

See axaff20828N002\_VV002\_vvref2.pdf for the full report

V&V Scientist	Beth Sundheim
V&V Date (YYYY-MM-DD)	2018.03.07
V&V Edition	2
V&V Disposition and Status	OK
V&V Charge Time	15.078400116086

## Comments

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One optional chip was dropped.
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A spatial region of the original bias map for CCD = 1 suffered from
anomalously high data values. Pixels in the event data that were
bias-corrected by one of the original affected bias pixels may have an
apparent energy shift. While the change in energy is expected to be
small (~20 eV), it depends on many parameters that have not yet been
fully explored for this bias anomaly. The bias map for CCD = 1 has been
reconstructed for this processing to remove this anomaly using scaled
data from a comparable bias map from another observation. The pixels
affected by the anomaly are bounded by sky coords: (47.83487,39.23896),
(47.83724,39.24116), (47.78510,39.27492), (47.78030,39.27430).
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The focal plane temperature during part of this observation was warmer
than the upper limit for optimum calibration of the ACIS gain and
spectral resolution (i.e., -114.0 C for ACIS-I and -112.0 C for
ACIS-S).
The Chandra calibration team calibrates the ACIS gain and spectral
resolution using data from the external calibration source (ECS). ECS
data show that the frontside-illuminated (FI) CCDs are more temperature
sensitive than the backside-illuminated (BI) CCDs.
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A summary of the current calibration status of the ACIS gain and
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spectral resolution can be found at:

http://asc.harvard.edu/cal/Acis/Cal\_prods/Gain\_and\_Spectral\_Resolution/A CIS\_response\_summary.html

The main points are: 1) The gain on BI chips remains within 0.3% (i.e., the systematic uncertainty in the ACIS gain quoted on the Chandra Calibration Status Summary web page) at all measured temperatures. 2) The gain on FI chips remains within 0.3% below row 600 at all measured temperatures. 3) The gain on FI chips above row 600 can be underestimated by as much as 1% for focal plane temperatures exceeding -116 C. 4) The spectral resolution (i.e., FWHM) on BI chips is insensitive to the focal plane temperature. 5) Warmer focal plane temperatures increase the FWHM on FI chips by up to 30 eV near row 512 and by up to 70 eV near the top of the chips. In summary, the user should be cautious in the spectral analysis of high S/N emission lines detected on the top half of FI chips in this observation. Default processing with the current version of the CALDB will underestimate photon energies by up to 1% and broaden emission lines by up to 70 eV.

seq_num	801801	Sequence number
obs_id	20828	Observation id
title	A Unique Sample of Extreme-BCG Clusters at $0.2 < z < 0.5$	Proposal
observer	Gordon Garmire	Principal investigator
object	SDSS395	Source name
dtycycle	0	
cycle	Р	events from which exps? Prim/Second/Both
ra_targ	47.602083	Observer's specified target RA [deg]
dec_targ	39.184361	Observer's specified target Dec [deg]
ra_nom	47.562133246717	Nominal RA [deg]
dec_nom	39.204690564423	Nominal Dec [deg]
roll_nom	140.2339093821	Nominal Roll [deg]
revision	2	Processing version of data
ontime	15078.400116086	Sum of GTIs [s]
livetime	14881.389718013	Livetime [s]
ontime0	15075.259025812	Sum of GTIs [s]
ontime1	15078.400116086	Sum of GTIs [s]
ontime2	15078.400116086	Sum of GTIs [s]
ontime3	15078.400116086	Sum of GTIs [s]
12events	39899	Number of level 2 events

