## V&V Summary Report L2 ASCDS Version: 10.6

Observation 20719 - L2 Version 1 Chandra X-Ray Center

L2 Processing Date: Feb 25 2018

See axaff20719N001\_VV001\_vvref2.pdf for the full report

V&V Scientist	Joy Nichols
V&V Date (YYYY-MM-DD)	2018.02.26
V&V Edition	1
V&V Disposition and Status	OK
V&V Charge Time	22.075900478601

## Comments

Big dither pattern used. z amplitude approximately 9 times y amplitude, with modified frequency. See archive metadata for details.

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Zeroth order is extended and piled up. The zeroth order sky position was determined using a software tool developed by CXC called findzero, which is available in CIAO as part of the tgdetect2 tool. The tool calculates the point of intersection of the readout streak on the ACIS CCD and the meg dispersed spectral arm, rather than using a centroid position of the source. The findzero results are more accurate than source centroid in this case.

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Faint grating spectra can be seen in an image of bad events. This is probably due to pileup in the spectrum, causing migration to bad grades. This should be considered in analysis.

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WARNING: there are no standard ciao tools for analysis of grating spectra from extended sources. The shape of an emission 'line' will be the shape of the zero order spatial structure convolved with the instrumental LSF. Grating extractions can be used, but need to be combined with custom spatial-spectral analysis, since wavelength is multi-valued at any particular diffraction angle.

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WARNING: Zeroth order selected by pipeline tools is well-centered in the extended source but is not at the position(s) of brightest emission. The user may want to select a region or source of interest, then use software tools such as CIAO to specify the coordinates of the zeroth order source of interest before running the tools to resolve the dispersed events.

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Gain and CTI correction are not well calibrated on CCD\_ID 5 (ACIS-S1). Default order sorting can clip some regions, particularly longward of 30A (first order). User-specified custom processing parameters may be required in tg\_resolve\_events (osipfile=none, osort\_lo, osort\_hi ~0.5) though this can allow more zeroth order background at short wavelengths.

seq_num	790317	Sequence number
obs_id	20719	Observation id
title	AO-19 Big Dither Observation of Mkn421	Proposal title
observer	CXC Calibration	Principal investigator
object	MKN421	Source name
dtycycle	0	<b>&amp;</b> #160
cycle	P	events from which exps? Prim/Second/Both
ra_targ	166.113333	Observer's specified target RA [deg]
dec_targ	38.208806	Observer's specified target Dec [deg]
ra_nom	166.13291815438	Nominal RA [deg]
dec_nom	38.197562003739	Nominal Dec [deg]
roll_nom	158.14453873959	Nominal Roll [deg]
revision	1	Processing version of data
ontime	22075.900478601	Sum of GTIs [s]
livetime	21281.892419601	Livetime [s]
ontime5	22075.900478601	Sum of GTIs [s]
ontime6	22075.900478601	Sum of GTIs [s]
ontime7	22075.900478601	Sum of GTIs [s]
ontime8	22075.900478601	Sum of GTIs [s]
ontime9	22074.759438634	Sum of GTIs [s]
12events	468626	Number of level 2 events

