## V&V Summary Report L2 ASCDS Version: 8.4.5

Observation 1783 - L2 Version 6 Chandra X-Ray Center

L2 Processing Date: Aug 30 2012

See axaff01783N004\_VV002\_vvref2.pdf for the full report

V&V Scientist	Beth Sundheim
V&V Date (YYYY-MM-DD)	2018.03.05
V&V Edition	2
V&V Disposition and Status	OK
V&V Charge Time	7.868

## Comments

The fid light in slot 2 was removed from the aspect solution due to poor data quality. The aspect solution is improved by the removal of this slot from the solution.

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The focal plane temperature during part of this observation was warmer than the upper limit for optimum calibration of the ACIS gain and spectral resolution (i.e., -114.0 C for ACIS-I and -112.0 C for ACIS-S).

The Chandra calibration team calibrates the ACIS gain and spectral resolution using data from the external calibration source (ECS). ECS data show that the frontside-illuminated (FI) CCDs are more temperature sensitive than the backside-illuminated (BI) CCDs.

A summary of the current calibration status of the ACIS gain and spectral resolution can be found at:

http://asc.harvard.edu/cal/Acis/Cal\_prods/Gain\_and\_Spectral\_Resolution/ACIS\_response\_summary.html

## The main points are:

1) The gain on BI chips remains within 0.3% (i.e., the systematic uncertainty in the ACIS gain quoted on the Chandra Calibration Status

Summary web page) at all measured temperatures.

- 2) The gain on FI chips remains within 0.3% below row 600 at all measured temperatures.
- 3) The gain on FI chips above row 600 can be underestimated by as much as 1% for focal plane temperatures exceeding -116 C.
- 4) The spectral resolution (i.e., FWHM) on BI chips is insensitive to the focal plane temperature.
- 5) Warmer focal plane temperatures increase the FWHM on FI chips by up to 30 eV near row 512 and by up to 70 eV near the top of the chips. In summary, the user should be cautious in the spectral analysis of high S/N emission lines detected on the top half of FI chips in this observation. Default processing with the current version of the CALDB will underestimate photon energies by up to 1% and broaden emission lines by up to 70 eV.

	1	
seq_num	590209	Sequence number
obs_id	1783	Observation id
title	ACIS CHIP RESPONSE TO LINES WITH E=0.6-1.5 KEV	Proposal title
observer	Dr. CXC Calibration	Principal investigator
object	E0102-72.3 [Chip I3, T=110, Offsets=-5,2,0]	Source name
dtycycle	0	<b>&amp;</b> #160
cycle	P	events from which exps? Prim/Second/Both
ra_targ	16.01	Observer's specified target RA [deg]
dec_targ	-72.032028	Observer's specified target Dec [deg]
ra_nom	16.091037335626	Nominal RA [deg]
dec_nom	-71.929369044456	Nominal Dec [deg]
roll_nom	100.88552813933	Nominal Roll [deg]
revision	6	Processing version of data
ontime	7868.6816902608	Sum of GTIs [s]
livetime	7769.0437047474	Livetime [s]
ontime0	7868.5585702658	Sum of GTIs [s]
ontime1	7868.5996102691	Sum of GTIs [s]
ontime2	7868.6406502724	Sum of GTIs [s]
ontime3	7868.6816902608	Sum of GTIs [s]
ontime6	7868.7637702674	Sum of GTIs [s]
ontime7	7868.7227302641	Sum of GTIs [s]
12events	75528	Number of level 2 events

