## V&V Summary Report L2 ASCDS Version : 10.8

## Observation 22734 - L2 Version 1 Chandra X-Ray Center

L2 Processing Date : Aug 25 2019

See axaff22734N001\_VV002\_vvref2.pdf for the full report

V&V Scientist	Beth Sundheim
V&V Date (YYYY-MM-DD)	2020.04.02
V&V Edition	2
V&V Disposition and Status	OK
V&V Charge Time	35.064100269794

## Comments

The guide star in slot 5 was removed from the aspect solution due to poor data quality. The aspect solution is improved by the removal of this slot from the solution.

```
==
The focal plane temperature during part of this observation was warmer
than the upper limit for optimum calibration of the ACIS gain and
spectral resolution (i.e., -112.0 C for ACIS-I).
The Chandra calibration team calibrates the ACIS gain and spectral
resolution using data from the external calibration source (ECS). ECS
data show that the frontside-illuminated (FI) CCDs are more temperature
sensitive than the backside-illuminated (BI) CCDs.
A summary of the current calibration status of the ACIS gain and
spectral resolution can be found at:
http://asc.harvard.edu/cal/Acis/Cal_prods/Gain_and_Spectral_Resolution/A
CIS_response_summary.html
```

The main points are: 1) The gain on BI chips remains within 0.3% (i.e., the systematic uncertainty in the ACIS gain quoted on the Chandra Calibration Status Summary web page) at all measured temperatures. 2) The gain on FI chips remains within 0.3% below row 600 at all measured temperatures.
 3) The gain on FI chips above row 600 can be underestimated by as much as 1% for focal plane temperatures exceeding -116 C.
 4) The spectral resolution (i.e., FWHM) on BI chips is insensitive to the focal plane temperature.
 5) Warmer focal plane temperatures increase the FWHM on FI chips by up to 30 eV near row 512 and by up to 70 eV near the top of the chips. In summary, the user should be cautious in the spectral analysis of high S/N emission lines detected on the top half of FI chips in this observation. Default processing with the current version of the CALDB will underestimate photon energies by up to 1% and broaden emission lines by up to 70 eV.

seq_num	801762	Sequence number
obs_id	22734	Observation id
title	Shock structure, the electron-ion equilibration timescale and the disintegrating cool core in A2146	Proposal title
observer	Helen Russell	Principal investigator
object	Abell 2146	Source name
dtycycle	0	
cycle	Р	events from which exps? Prim/Second/Both
ra_targ	239.06125	Observer's specified target RA [deg]
dec_targ	66.346917	Observer's specified target Dec [deg]
ra_nom	238.98781313544	Nominal RA [deg]
dec_nom	66.313884049762	Nominal Dec [deg]
roll_nom	274.27594145606	Nominal Roll [deg]
revision	1	Processing version of data
ontime	35064.100269794	Sum of GTIs [s]
livetime	34605.961985955	Livetime [s]
ontime0	35051.5362885	Sum of GTIs [s]
ontime1	35045.254208207	Sum of GTIs [s]
ontime2	35060.959289551	Sum of GTIs [s]
ontime3	35064.100269794	Sum of GTIs [s]
ontime6	35060.959249616	Sum of GTIs [s]
12events	143896	Number of level 2 events

