V&V Summary Report L2 ASCDS Version: 7.6.9

Observation 1046 - L2 Version 001 Chandra X-Ray Center

L2 Processing Date: Nov 29 2006

See axaff01046N001_VV001_vvref2.pdf for the full report

V&V Scientist	David Huenemoerder
V&V Date (YYYY-MM-DD)	2006.12.04
V&V Edition	1
V&V Disposition and Status	OK
V&V Charge Time	69.932

Comments

Roll constraint met.

WARNING: there are no standard ciao tools for analysis of grating spectra

from extended sources. The shape of an emission 'line' will be the shape of the zero order spatial structure convolved with the instrumental LSF. Grating extractions can be used, but need to be combined with custom spatial-spectral analysis, since wavelength is multi-valued at any particular diffraction angle.

WARNING::Zeroth order selected by pipeline tools is on a bright filament southeast of the center of the supernova remnant. The user will need to select a region or source of interest, then use software tools such as CIAO to specify the coordinates of the zeroth order source of interest before running the tools to resolve the dispersed events. The spectral data supplied in this processing are only energy-calibrated for the emission knot R1 (Lazendic et al. 2006, ApJ, 651, 250). However, it should be noted that the emission knot R1 that has been selected as the zeroth order source is filamentary and curved, so the energy assignments to the events should take the spatial information into account. The zeroth order used for extracting the spectral data in this

processing is not located at the position of the brightest X-ray emission in the filament, but closer to the inner curve of the filament.

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HIGH RESOLUTION SPECTRA OF YOUNG SUPERNOVA REMNANTS
Prof. Claude Canizares
CAS A
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P
350.866667
58.811806
350.87057410749
58.815806203148
85.955941473876
2
69942.277184382
69056.62595649
69942.277363941
69952.000065148
69939.036264181
69942.277184382
69942.277254254
69945.51816465
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